

Bringing Light into the Darkness

— *About Dark Matter Scattering off Light Nuclei*

In Collaboration with **Jordy de Vries** and **Andreas Nogga**

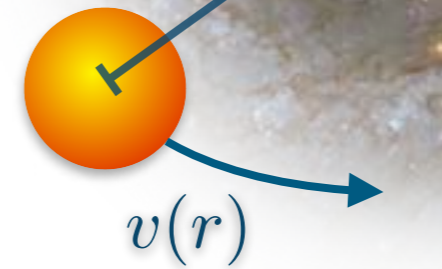
[arxiv:1704.01150]

13. July 2016 | **Christopher Körber** | IKP-3/IAS-4 | Seminar@TP2 — Ruhr-Universität Bochum

PhD Advisor: **Tom Luu**

Dark Matter

– So far only Gravitational Evidence $r > 8.5$ kpc



Rotation Curves: Milky Way

- Newtonian prediction

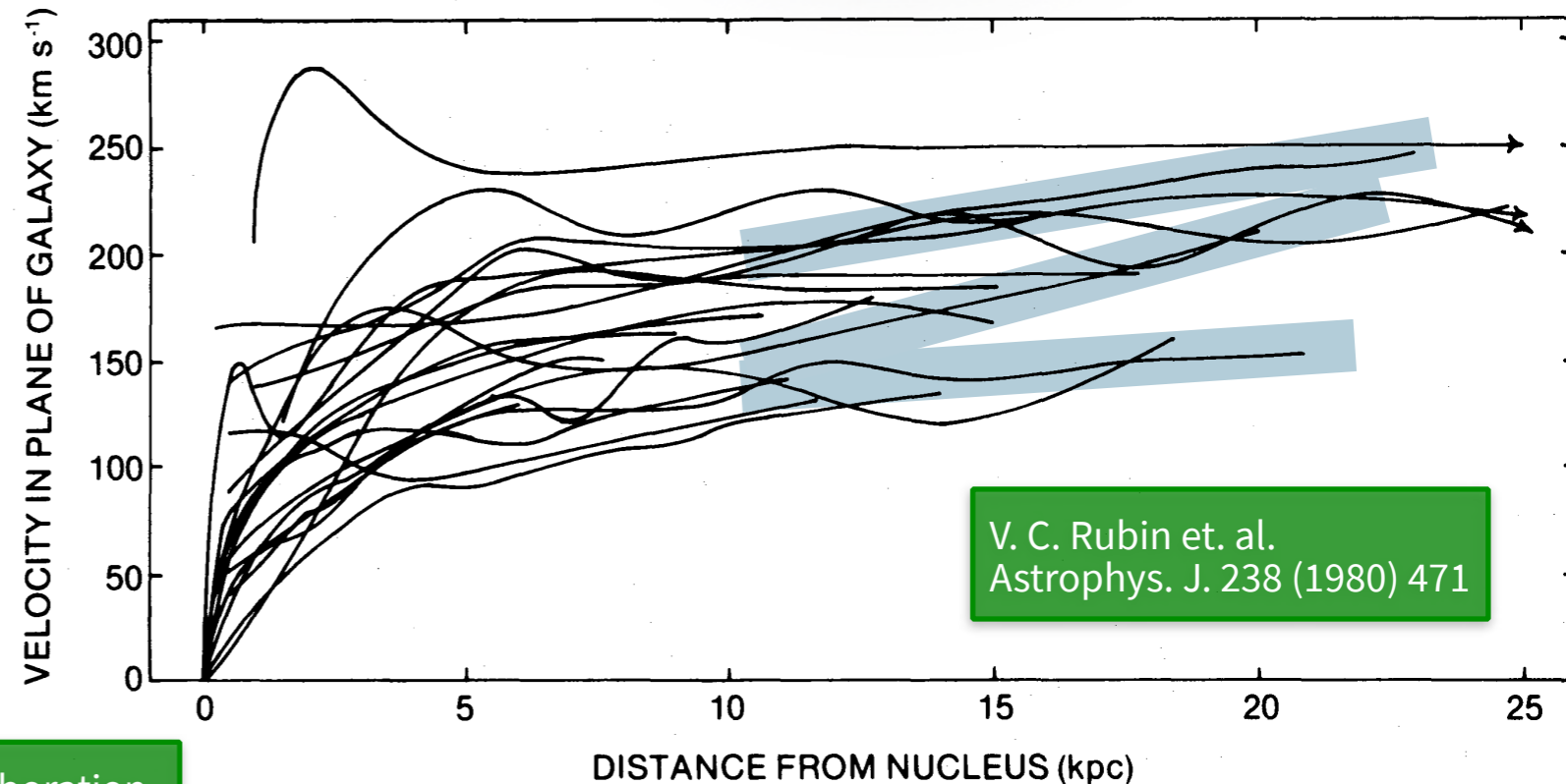
$$v(r) \propto \frac{1}{\sqrt{r}}$$

- Experimental observations

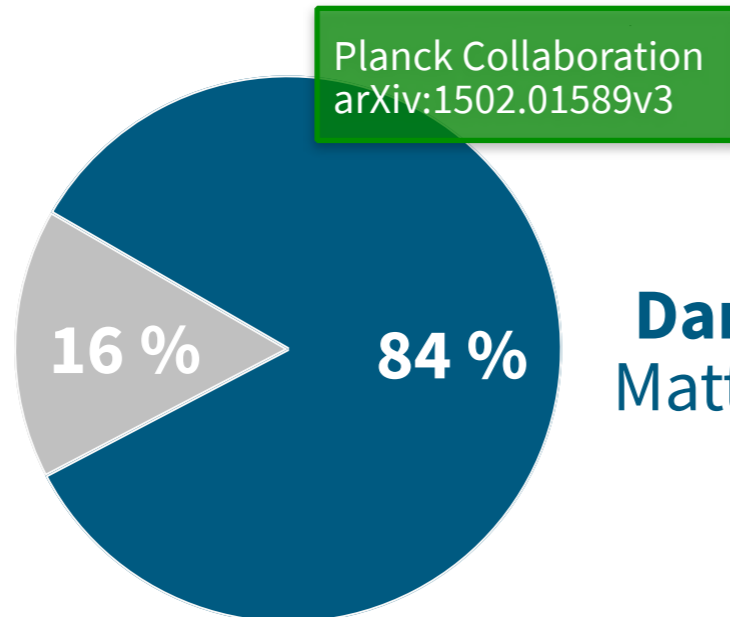
$$v(r) \propto r \Rightarrow \frac{M_{DM}}{M_M} \approx 20$$

Cosmological Parameters

- Λ CDM model



Baryonic
Matter

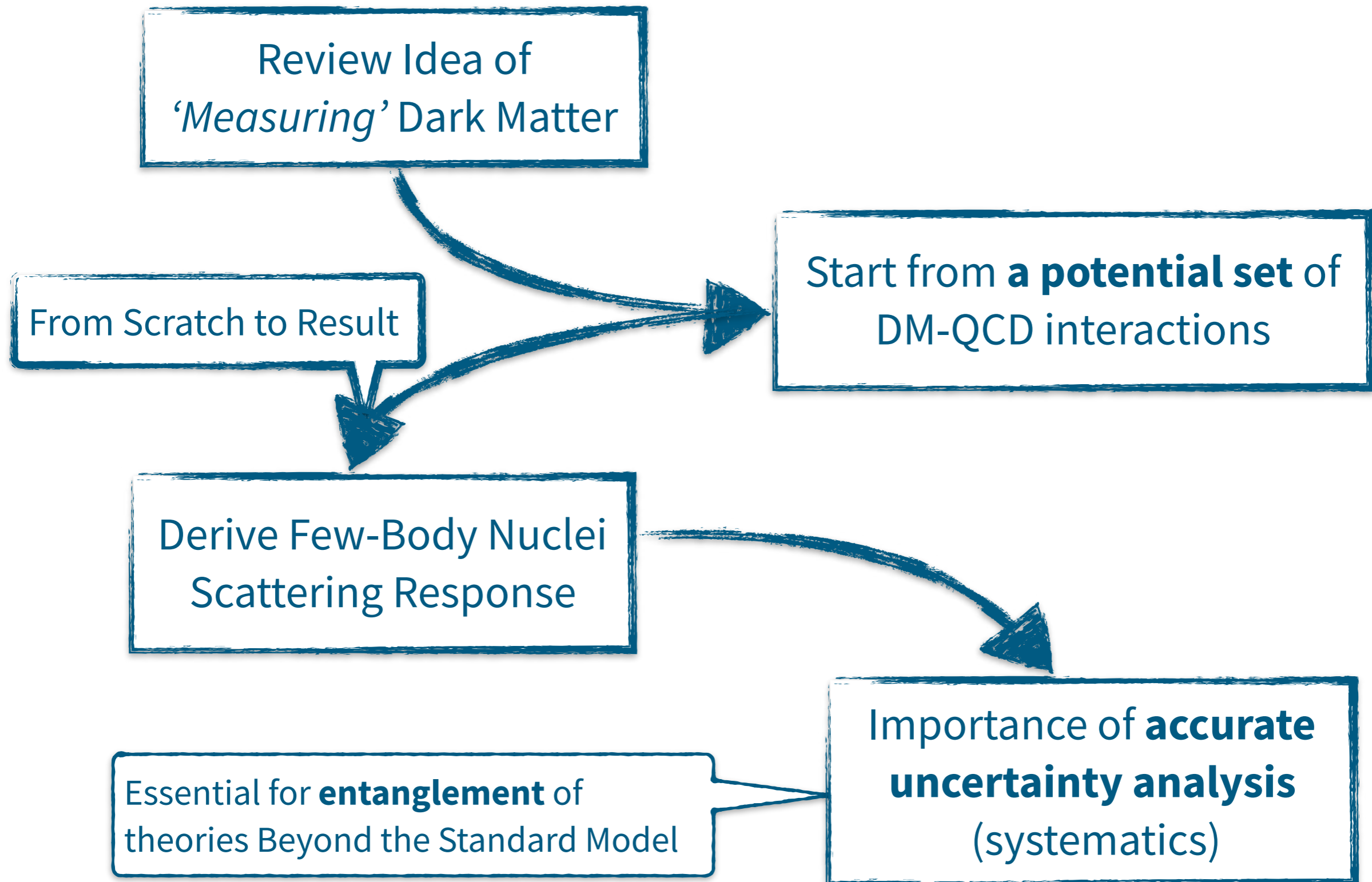


Dark
Matter

How to resolve this?

- **Assumption:** Existence of DM particle candidate
- **Connect experiments** on fundamental scales to theories **Beyond the Standard Model**

The Next ~40 Minutes



DM Annihilation

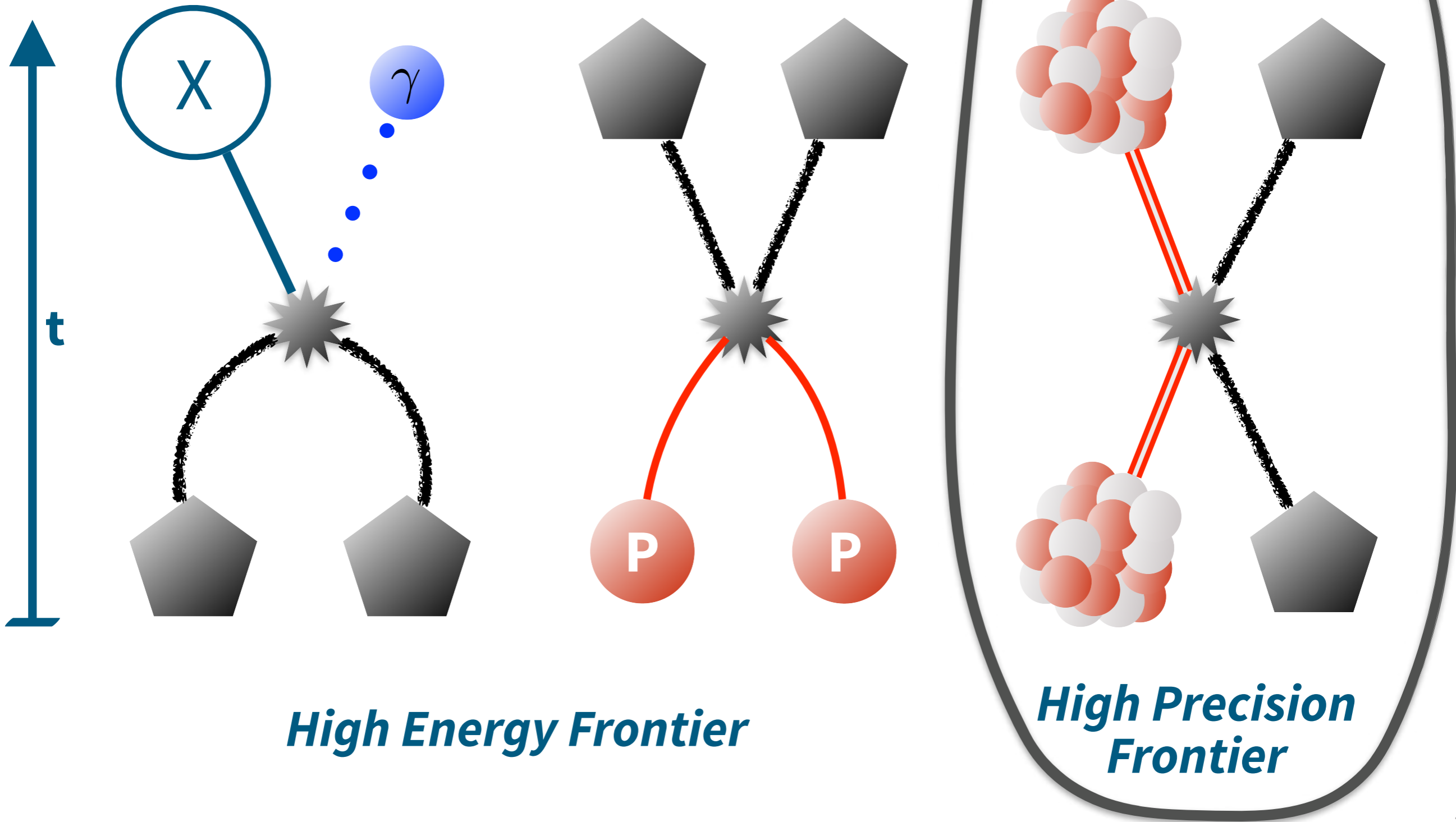
Astronomical Observation

DM Creation

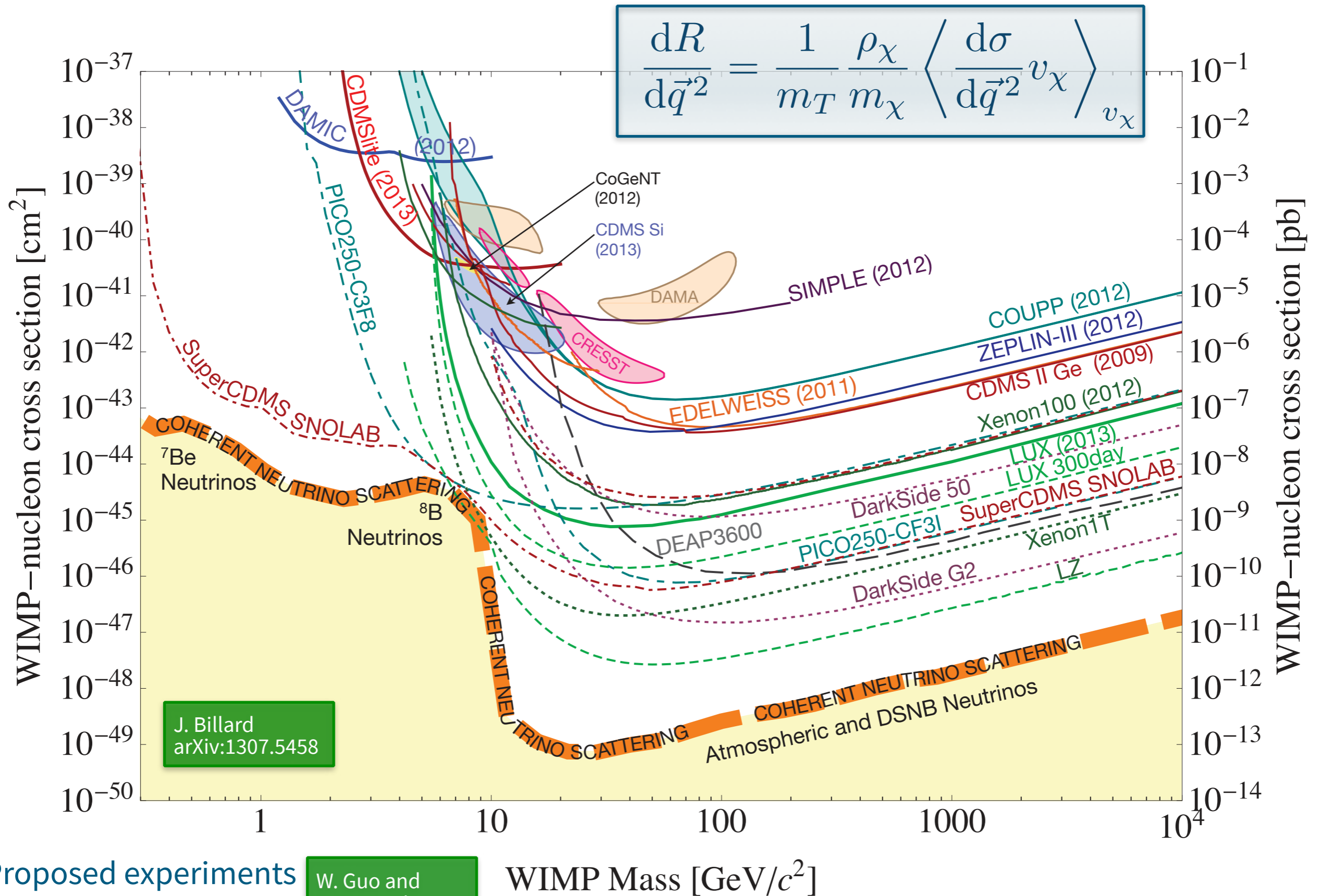
Collider Experiments

DM Scattering

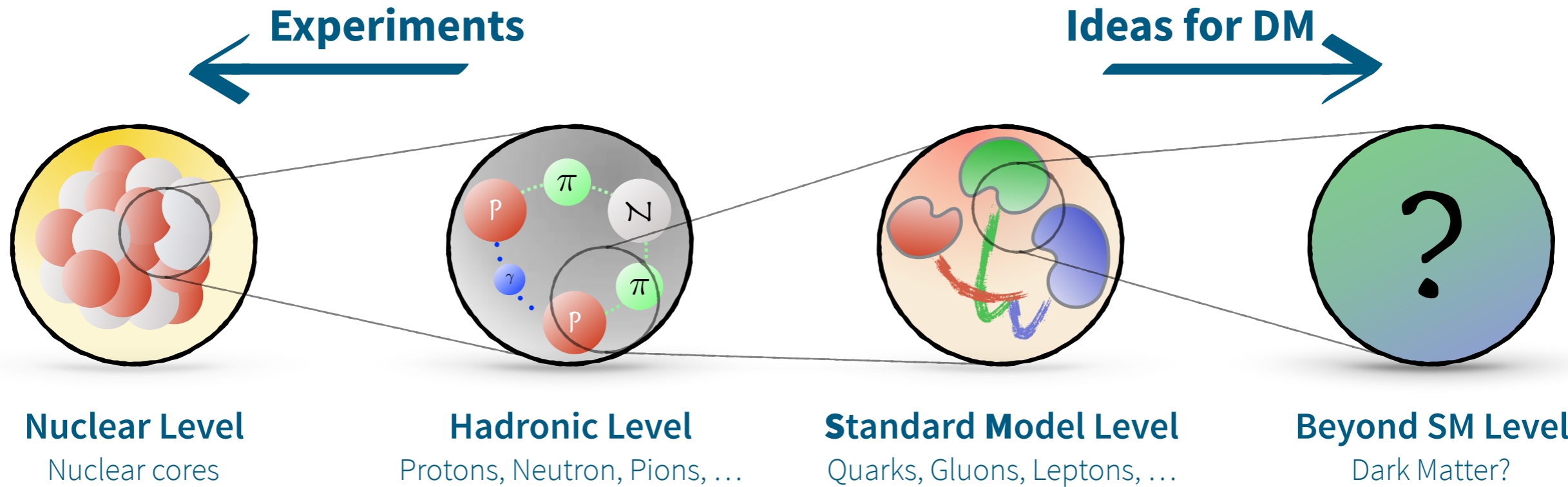
Underground Experiments



Direct Detection Bounds

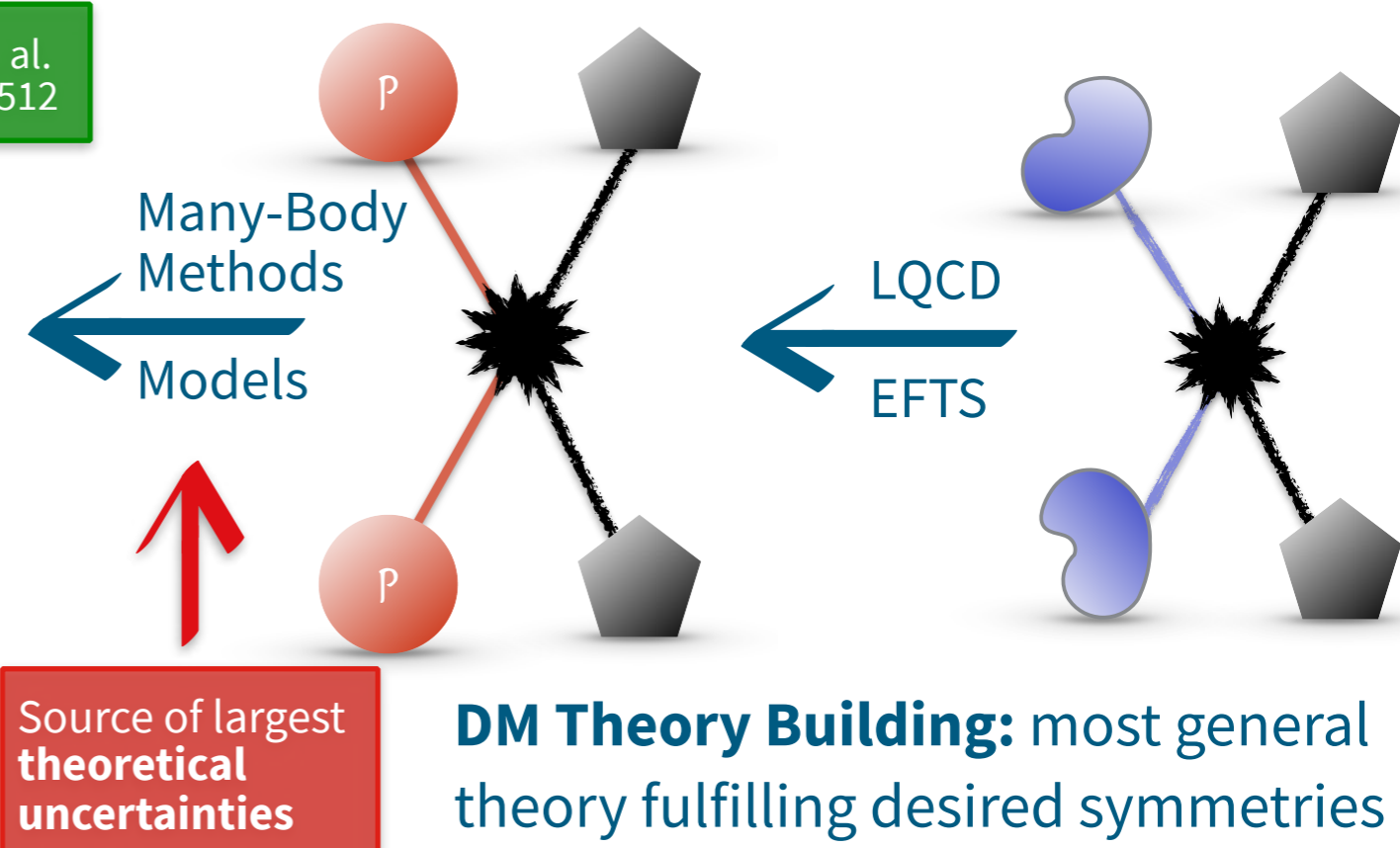
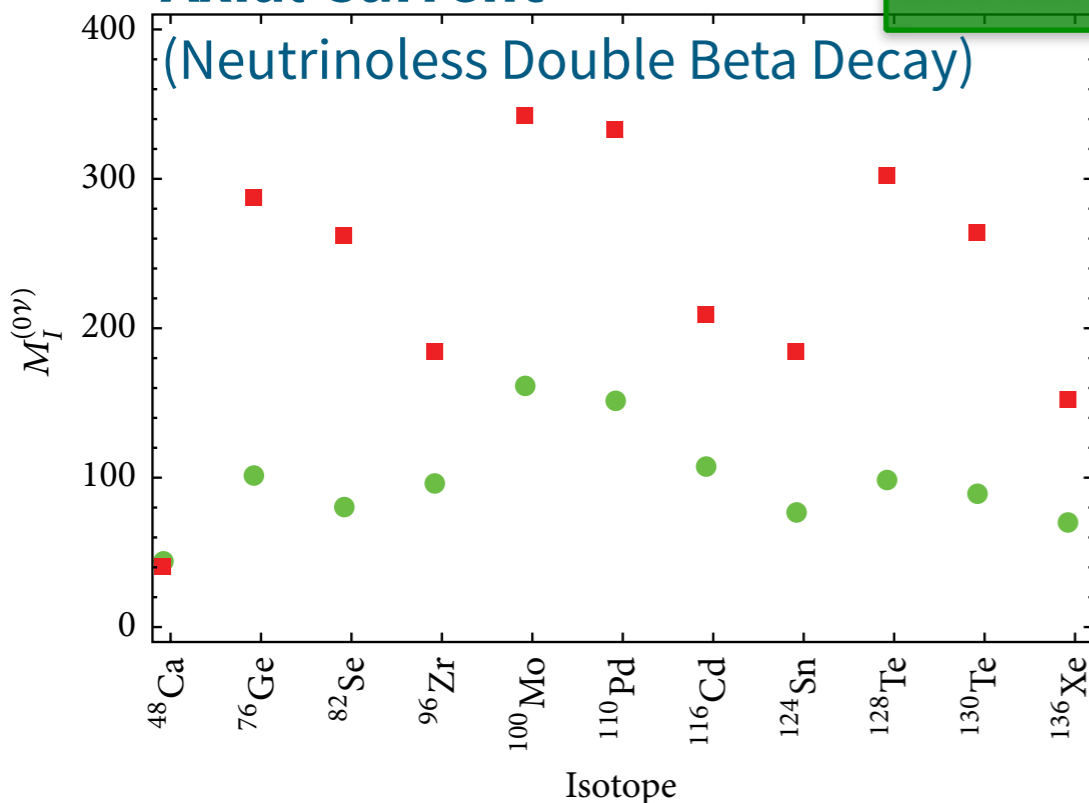


From Theory to Experiment



Nuclear Matrix Element Axial Current

S. Dell'Oro et. al.
arXiv:1601.07512



A Dark-Matter EFT

Dark Matter Interactions

– From Quarks to Nucleons

Effective Approaches

- M. Hoferichter, A. Schwenk, ...
arXiv:1605.08043
- F. L. Fitzpatrick, W. Haxton, ...
arXiv:1203.3542
- V. Cirigliano, M. L. Graesser, ...
arXiv:1205.2695

Possible Dark Matter Interaction

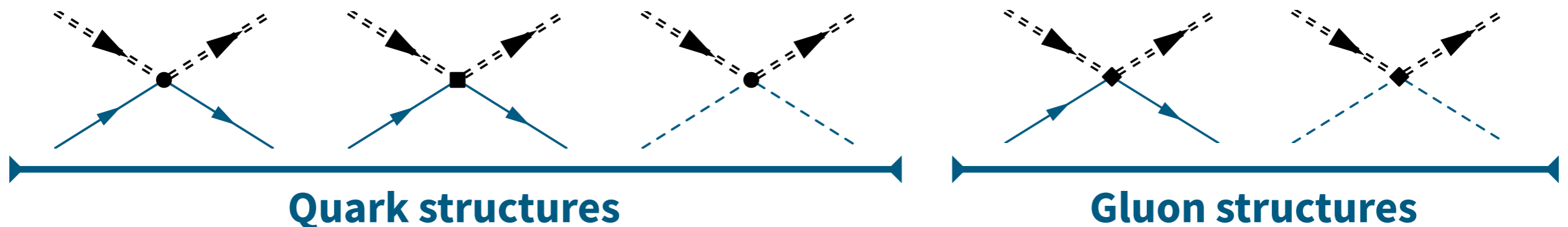
- Scalar interaction of DM with quarks and gluons

$$\mathcal{L}_{DM}^{(QCD)} = \bar{\chi}\chi \left[\sum_{f=u,d,s} \bar{c}_f m_f \bar{q}_f q_f + \bar{c}_g \alpha_s G_{\mu\nu}^a G^{\mu\nu a} \right]$$

Propagation of Scales

- Effective chiral Lagrangian based on underlying interactions
- Depends on chiral as well as lattice input (sigma term, mass splitting, ...)
- Gluon and strange quark structures can be grouped together (similar terms)

$$\mathcal{L}_{DM}^{(N\pi)} = \bar{\chi}\chi \left[c_q^{(ic)} \bar{N}N + c_q^{(iv)} \bar{N}\tau^3 N + c_q^{(\pi)} \pi^2 + c_g^{(ic)} \bar{N}N + c_g^{(\pi)} \left((\partial_\mu \pi)^2 - 2m_\pi^2 \pi^2 \right) \right]$$



Dark Matter Interactions

– From QCD to Chiral Lagrangian

Idea: Compare to existing framework of regular Chiral Perturbation Theory (ChPT)

A Primer for Chiral Perturbation Theory
S. Scherer and M. R. Schindler

$$\mathcal{L}_{DM}^{(QCD)} = (\bar{c}_u m_u \bar{\chi}\chi) \bar{u}u + (\bar{c}_d m_d \bar{\chi}\chi) \bar{d}d + \dots$$

- DM interaction as mass term $\chi_{\mathcal{M}} = 2B_0 \mathcal{M} \mapsto 2B_0 [\mathcal{M} - \text{diag}(\bar{c}_u m_u \bar{\chi}\chi, \bar{c}_d m_d \bar{\chi}\chi)]$

$$\mathcal{L}^{(N\pi)} \mapsto \mathcal{L}_{DM}^{(N\pi)} = \frac{F_0^2}{4} \text{Tr} [U^\dagger \chi_{\mathcal{M}} + U \chi_{\mathcal{M}}^\dagger] + c_1 \text{Tr} [\chi_{\mathcal{M}}^\dagger] \bar{N}N + \dots$$

- Expand interaction $U = u^2 = e^{i\pi \cdot \tau / F_0}$ $\chi_{\mathcal{M}}^\dagger = u^\dagger \chi_{\mathcal{M}} u^\dagger + h.c.$
- Iso-scalar interaction gets correction from one loop

$$\mathcal{L}_{DM}^{(N\pi)} = \bar{\chi}\chi [c_q^{(\pi)} \boldsymbol{\pi}^2 + c_q^{(ic)} \bar{N}N] + \dots$$

$$c_q^{(ic)} = \frac{\sigma_{\pi N}}{2} \bar{c}_{u,d} \quad c_q^{(\pi)} = \frac{m_\pi^2}{4} \bar{c}_{u,d} \quad \bar{c}_{u,d} = \bar{c}_u(1 - \epsilon) + \bar{c}_d(1 + \epsilon) \quad \epsilon = \frac{m_d - m_u}{m_d + m_u} \simeq 0.37$$

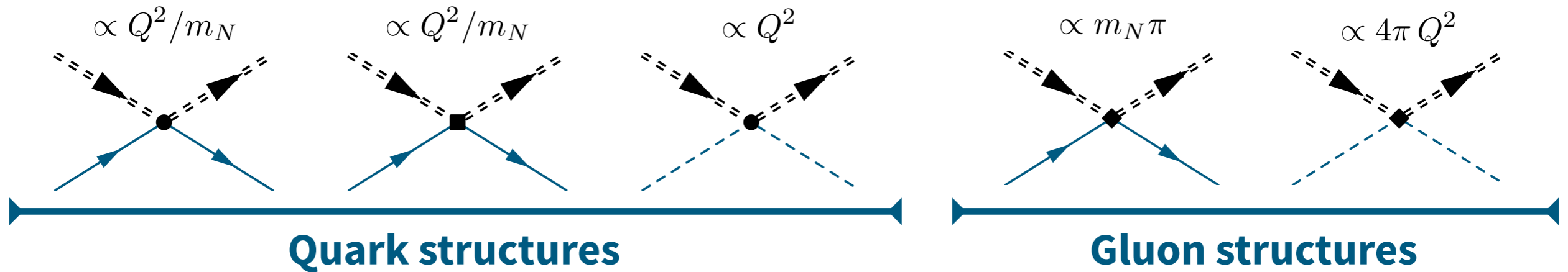
One-Body vs. Two-Body interactions

- DM **two-pion** exchange and **contact** interactions have **same ‘quark-scaling’**
- Size of **nucleon sigma term** indicates relative strength of **one-body vs. two-body**

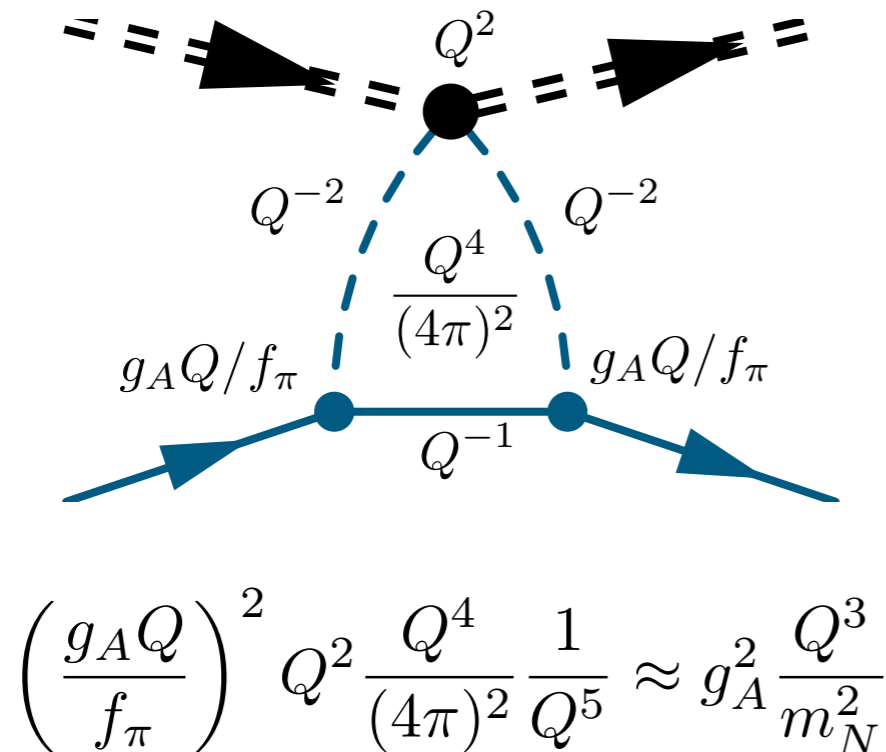
Power Counting

— Operator Counting

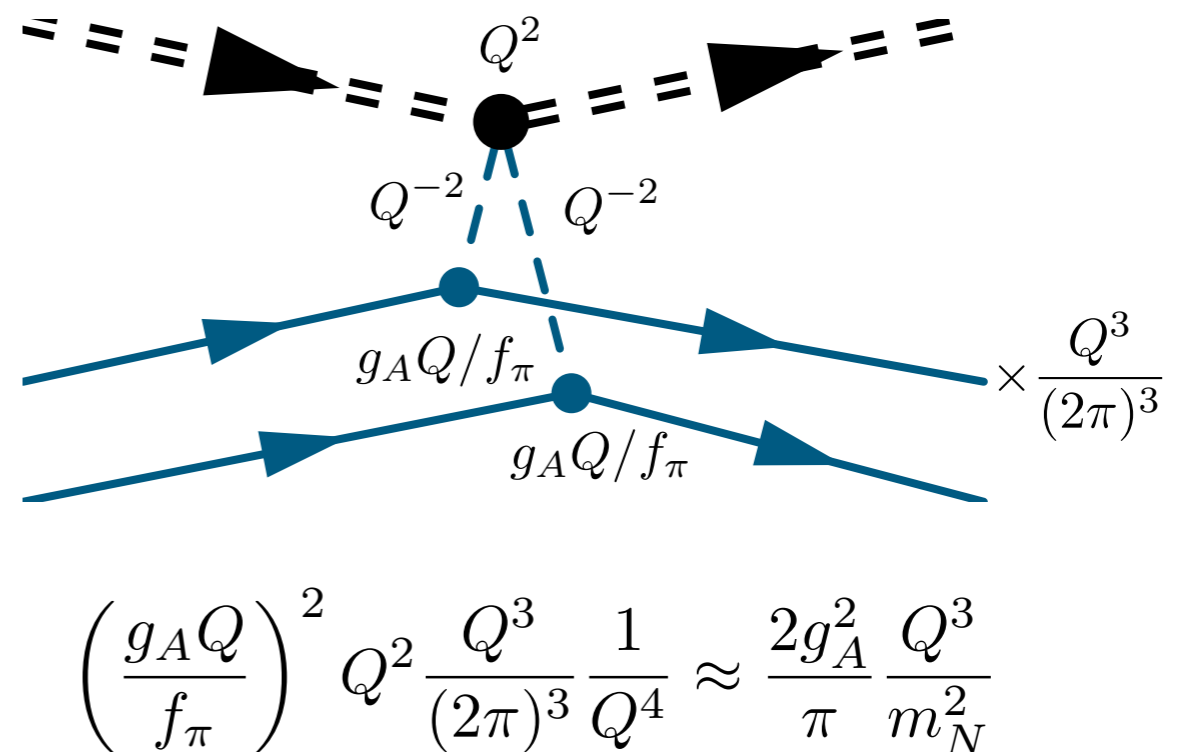
$$\propto g_A^2 Q^3 / m_N^2$$



Iso-Scalar Loop



Two-Pion Exchange



From Nucleons to Nuclei

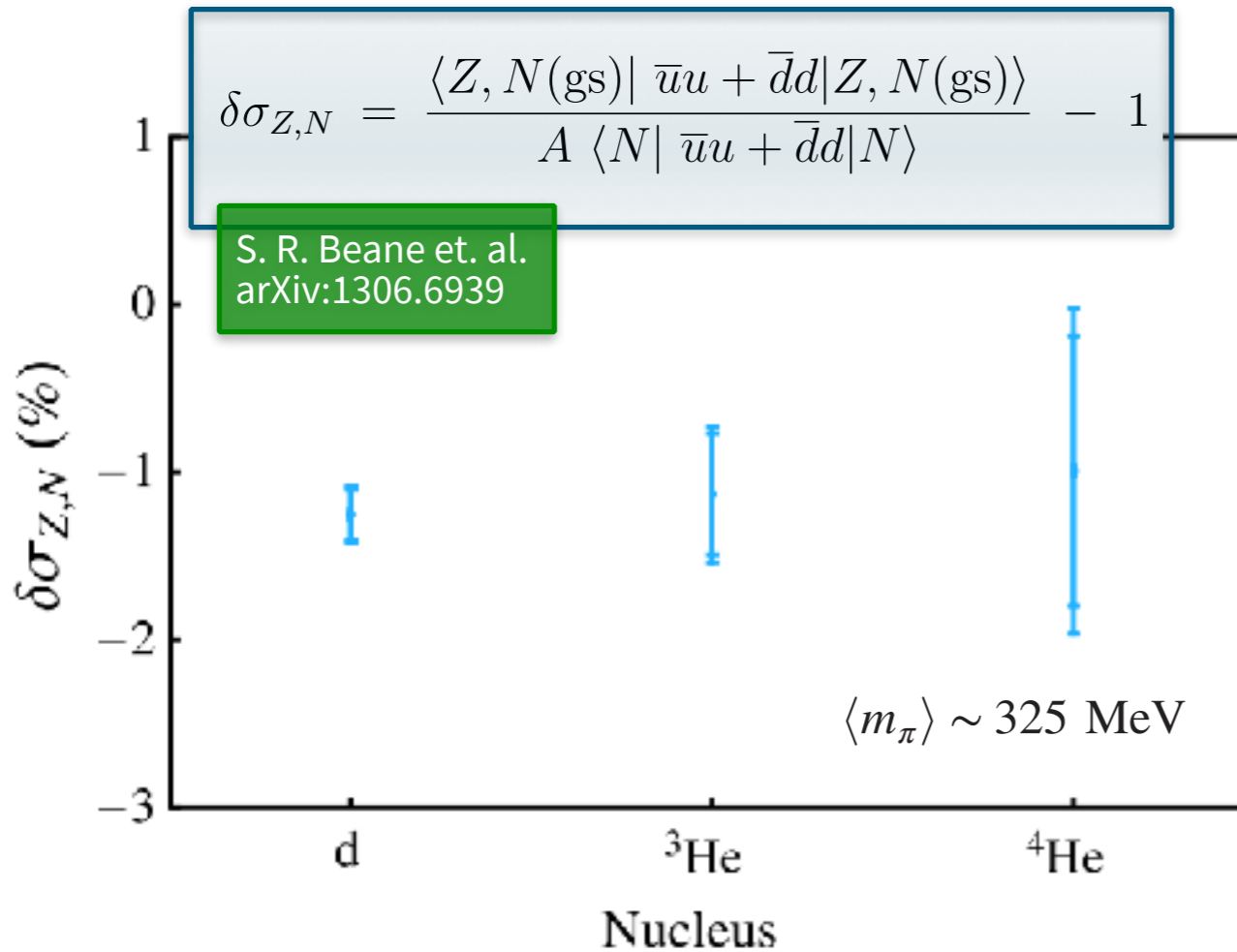
Dark Matter Interactions

— From Nucleons to Nuclei

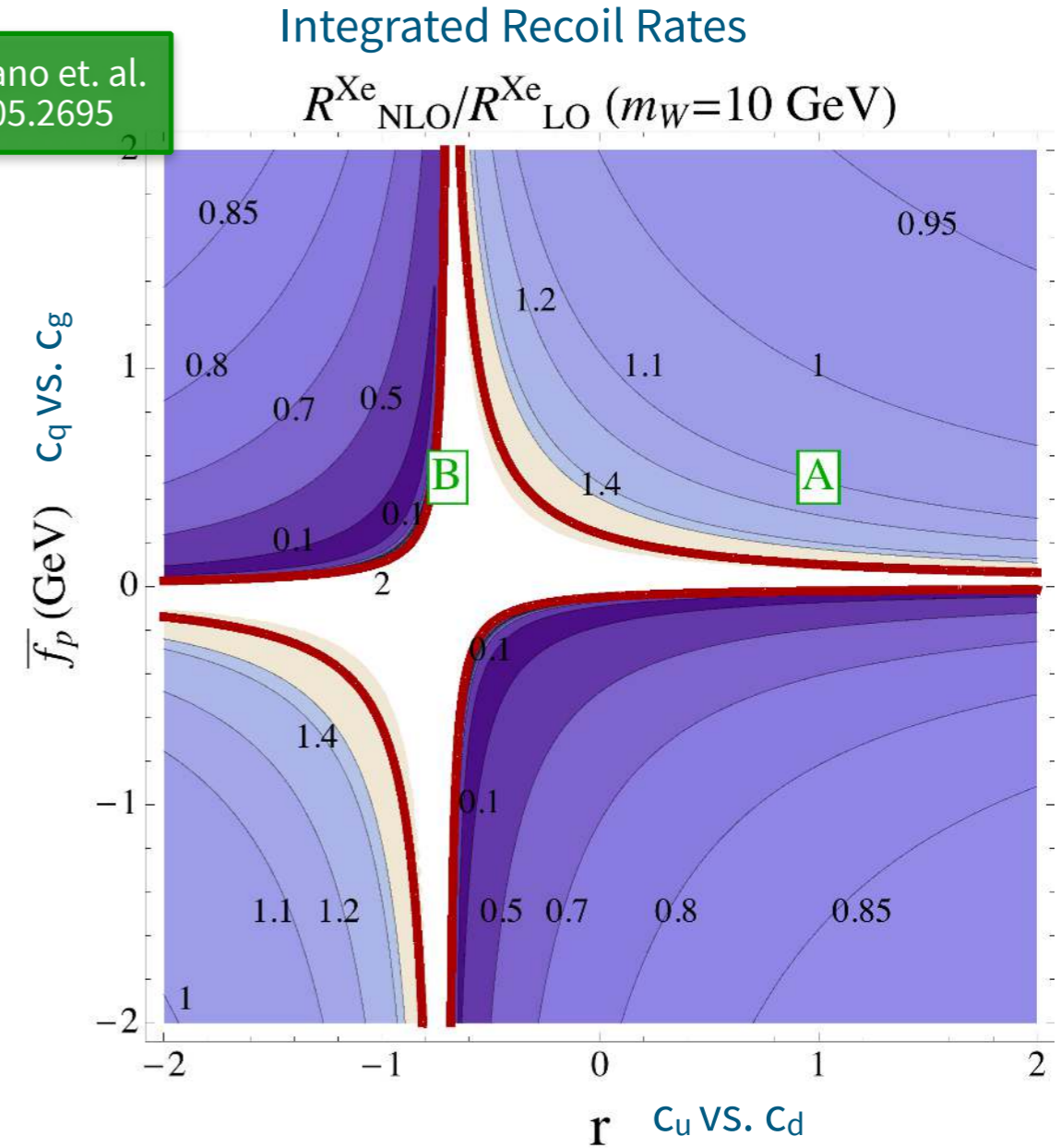
Challenges

- Possible cancellations of leading structures (*fine-tuning*)
- Relevance of two-body effects and meson exchange

Lattice Results



V. Cirigliano et. al. arXiv:1205.2695



⇒ need accurate and precise description

Light Nuclei Methodology

– Deterministic Scattering Solutions

Nuclear Forces:

- **Chiral Forces** — E. Epelbaum, H. Krebs, U. G. Meißner
arXiv:1412.4623
- **AV18** — R. B. Wiringa et. al.
arXiv:nucl-th/9408016
- **CDB** — R. Machleidt
arXiv:nucl-th/0006014

Obtain Nuclear Wave

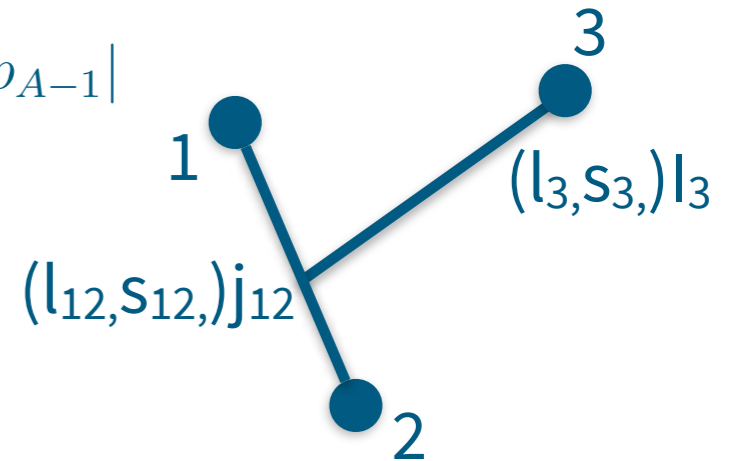
- Solve Schrödinger or Faddeev equation (^2H , ^3H , ^3He)

$$\hat{H} |\psi, \mathbf{P}, J, m_\tau\rangle = E |\psi, \mathbf{P}, J, m_\tau\rangle$$

$$\hat{\mathbb{1}} = \sum_\alpha \prod_{i=1}^{A-1} \int dp_i p_i^2 |\alpha, p_1, \dots, p_{A-1}\rangle \langle \alpha, p_1, \dots, p_{A-1}|$$

LSJ-Coupling scheme

$$|\alpha\rangle |_{A=3} := |((l_{12} s_{12}) j_{12} (l_3 s_3) I_3) j_3 m_{j_3}\rangle \otimes |(t_{12} t_3) \tau_3 m_{\tau_3}\rangle$$



Compute DM-Scattering Amplitude

- Describe DM-Interaction as momentum dependent current (perturbative)

$$\frac{d\sigma}{d\vec{q}^2} \propto \left| \sum_{\nu, t} \alpha_t \mathcal{F}_t^{(\nu)}(\vec{q}^2) \right|^2 \quad \mathcal{F}_t^{(\nu)}(\vec{q}^2) \propto \langle \psi, \vec{P} + \vec{q}, J, m_\tau | \hat{J}_t^{(\nu)}(\vec{q}) | \psi, \vec{P}, J, m_\tau \rangle$$

Leading response function normalized to 1 at $q = 0$ MeV

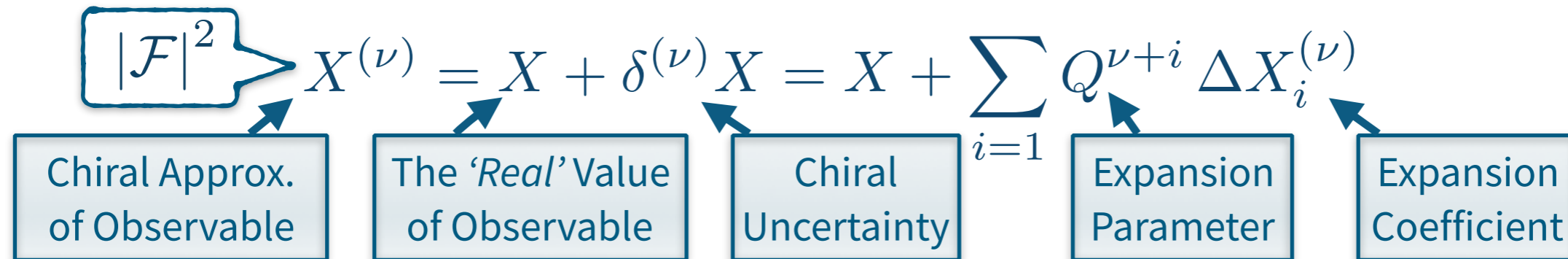
Relevant for light nuclei: $q \sim$ few MeV

Uncertainty of Results

- Convergence of basis (numerical uncertainty $< 1\%$)
- Nuclear uncertainty (analyze chiral behaviour)

Convergence of ChPT observables

- ChPT provides systematic expansion of observables in small parameter: $Q \sim \frac{m_\pi}{\Lambda}$



$$\nu < \nu' \Rightarrow X^\nu - X^{\nu'} = Q^{\nu+1} \left(\Delta X_1^{(\nu)} + \mathcal{O}(Q) \right)$$

The Estimate (very conservative)

- Estimate chiral uncertainty by approximating expansion coefficients

$$\delta^{(\nu)} X \leq Q^\nu \sum_{i=1}^{\infty} Q^i \left| \Delta X_i^{(\nu)} \right| \leq Q^\nu \left| \Delta X_{\max}^{(\nu)} \right| \sum_{i=1}^{\infty} Q^i \leq \frac{Q^{\nu+1}}{1-Q} \left| \Delta X_{\max} \right|$$

- Approximate maximal coefficient as maximum over chiral orders

$$\left| \Delta X_{\max}^{(\nu)} \right| \leq \left| \Delta X_{\max} \right| = \max_{\nu} \left(\left| \Delta X_{\max}^{(\nu)} \right| \right) = \max_{\nu} \left(\max_{\nu' > \nu} \left(\frac{\left| X^{(\nu)} - X^{(\nu')} \right|}{Q^{\nu+1}} \right) \right)$$

Our Results

Chiral Uncertainty Estimation

— Deuteron Quark-DM Amplitudes

cut-off	Λ_1	Λ_2	Λ_3	Λ_4	Λ_5
R	0.8 fm	0.9 fm	1.0 fm	1.1 fm	1.2 fm
Λ_b	600 MeV	600 MeV	600 MeV	500 MeV	400 MeV

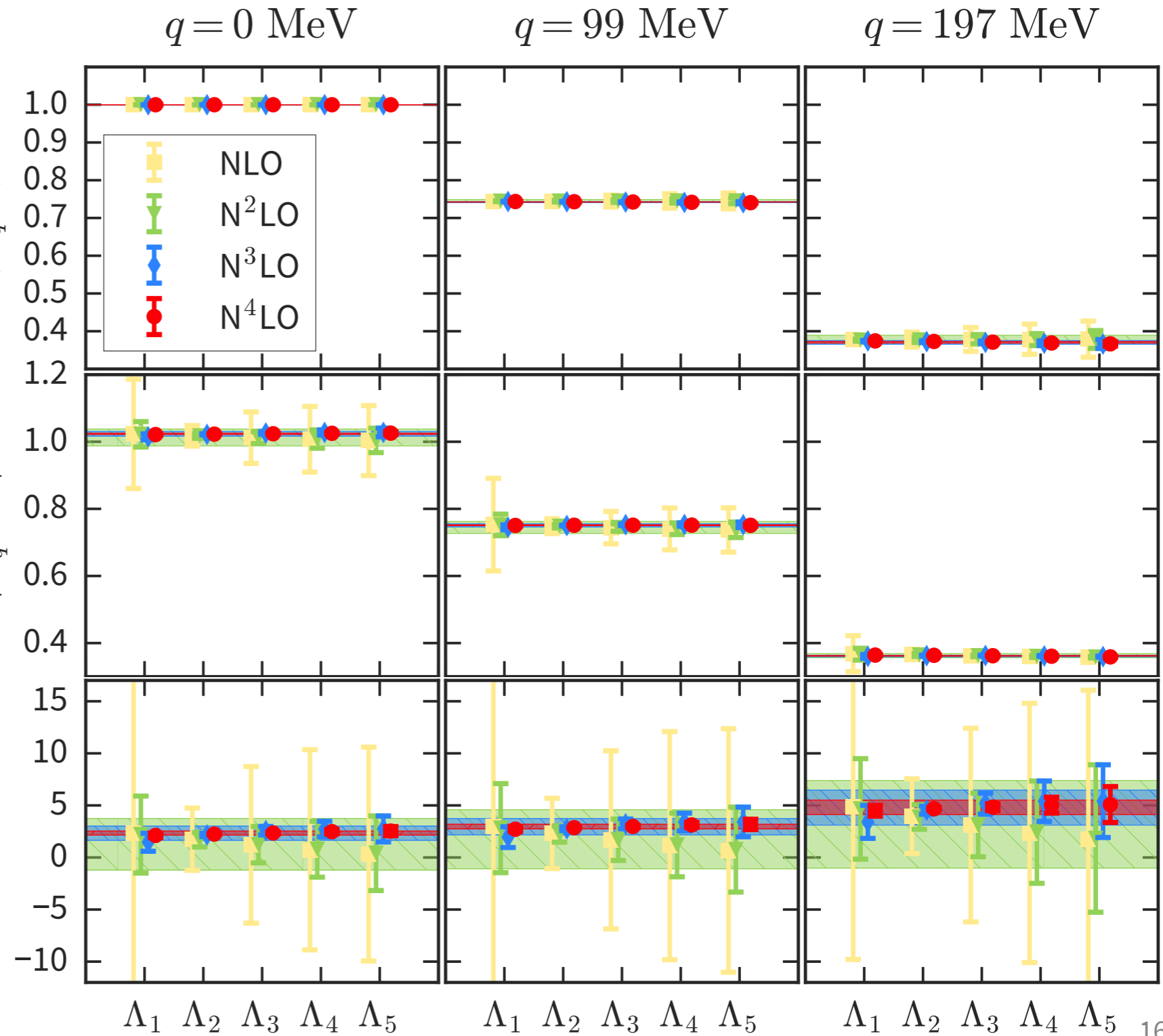
- Results for $c_u = c_d$
- Isoscalar DM interactions
- **Fixed order current**

Leading Order Currents

Next to Leading Order Currents

2-Body vs. 1-Body Currents

- Results **consistent** over different cut-offs
- Second cut-off gives best results



Chiral Uncertainty Estimation

— Helion Quark-DM Amplitudes

cut-off	Λ_1	Λ_2	Λ_3	Λ_4	Λ_5
R	0.8 fm	0.9 fm	1.0 fm	1.1 fm	1.2 fm
Λ_b	600 MeV	600 MeV	600 MeV	500 MeV	400 MeV

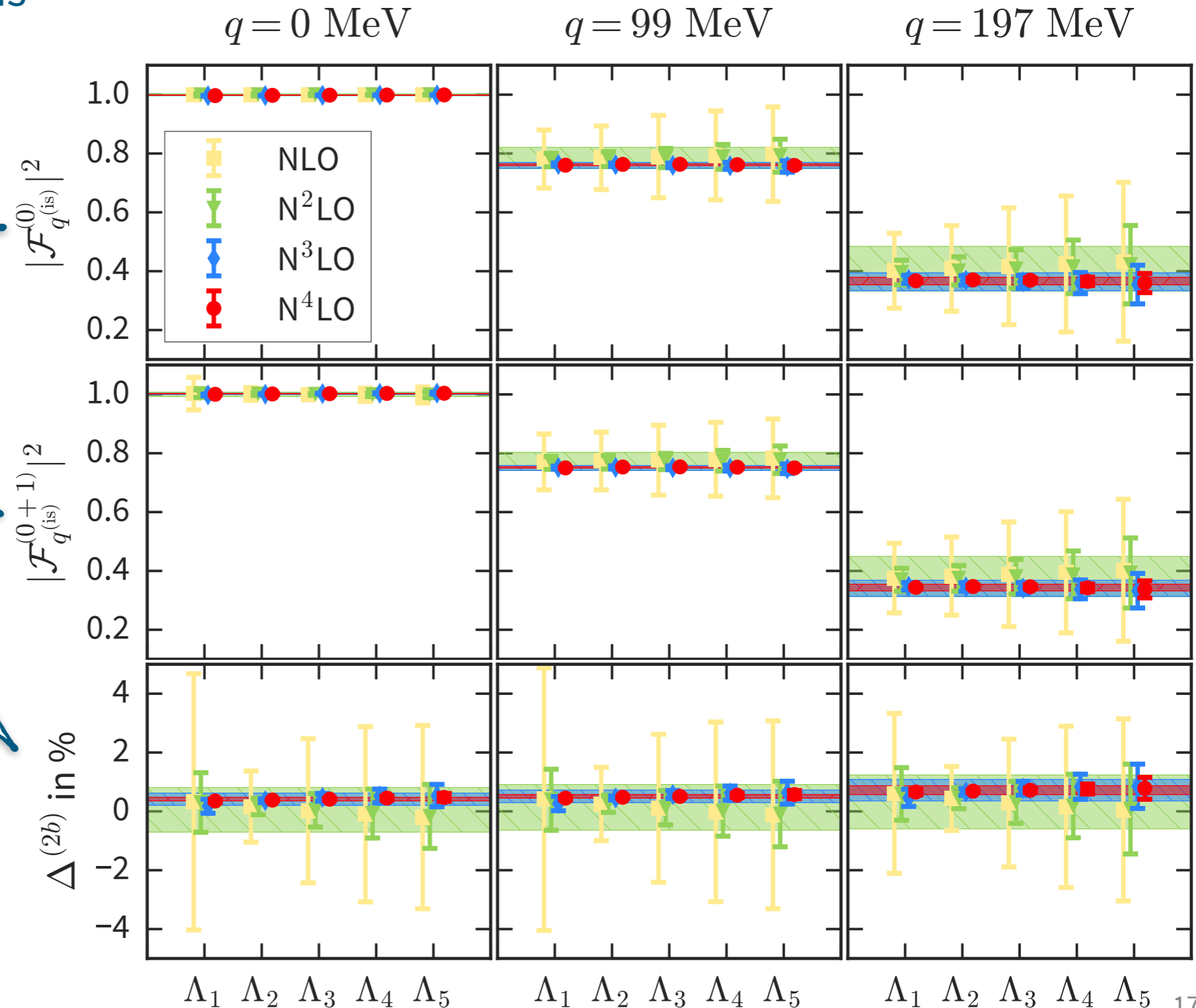
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Leading Order
Currents

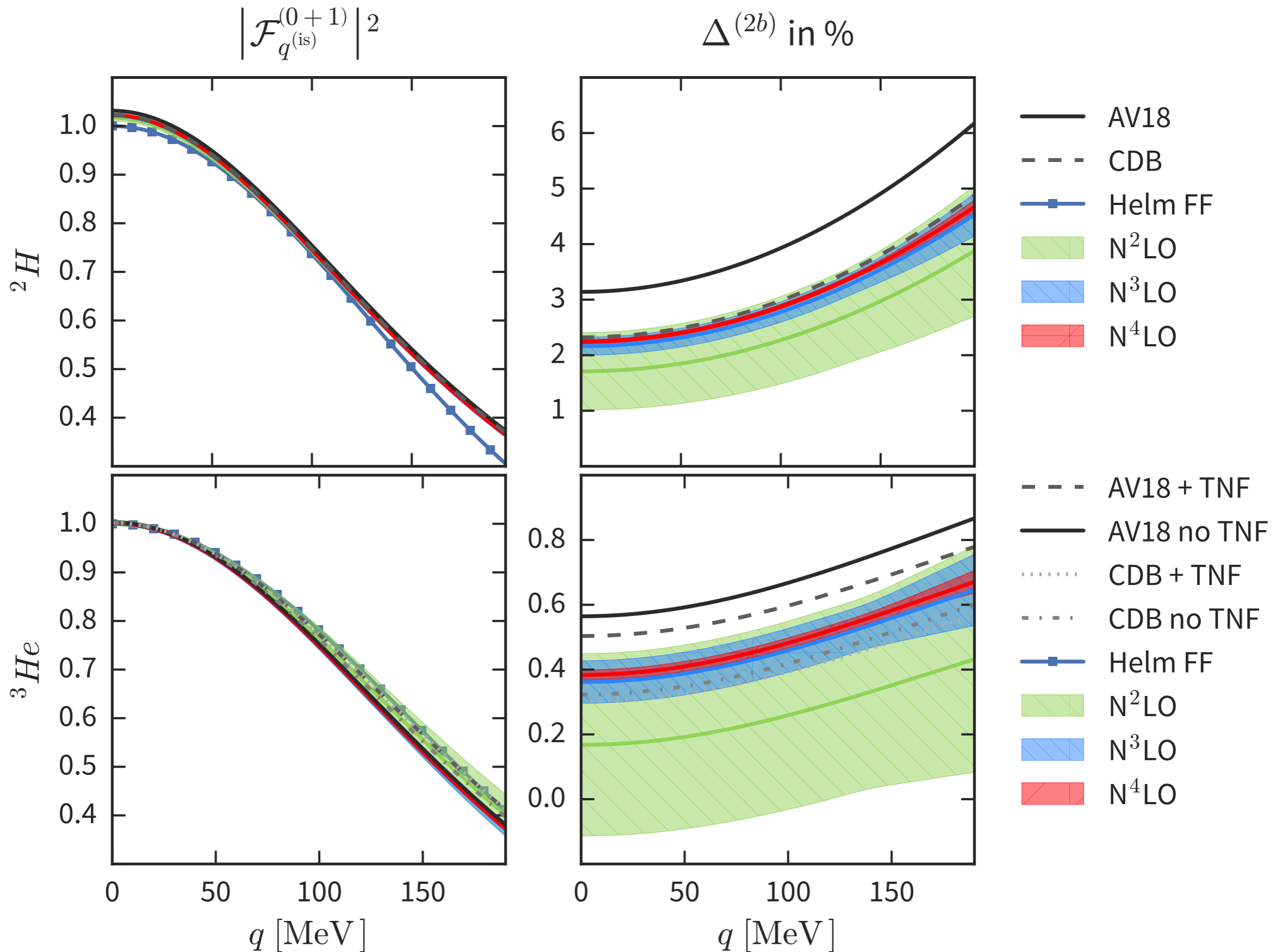
Next to Leading
Order Currents

2-Body vs. 1-Body
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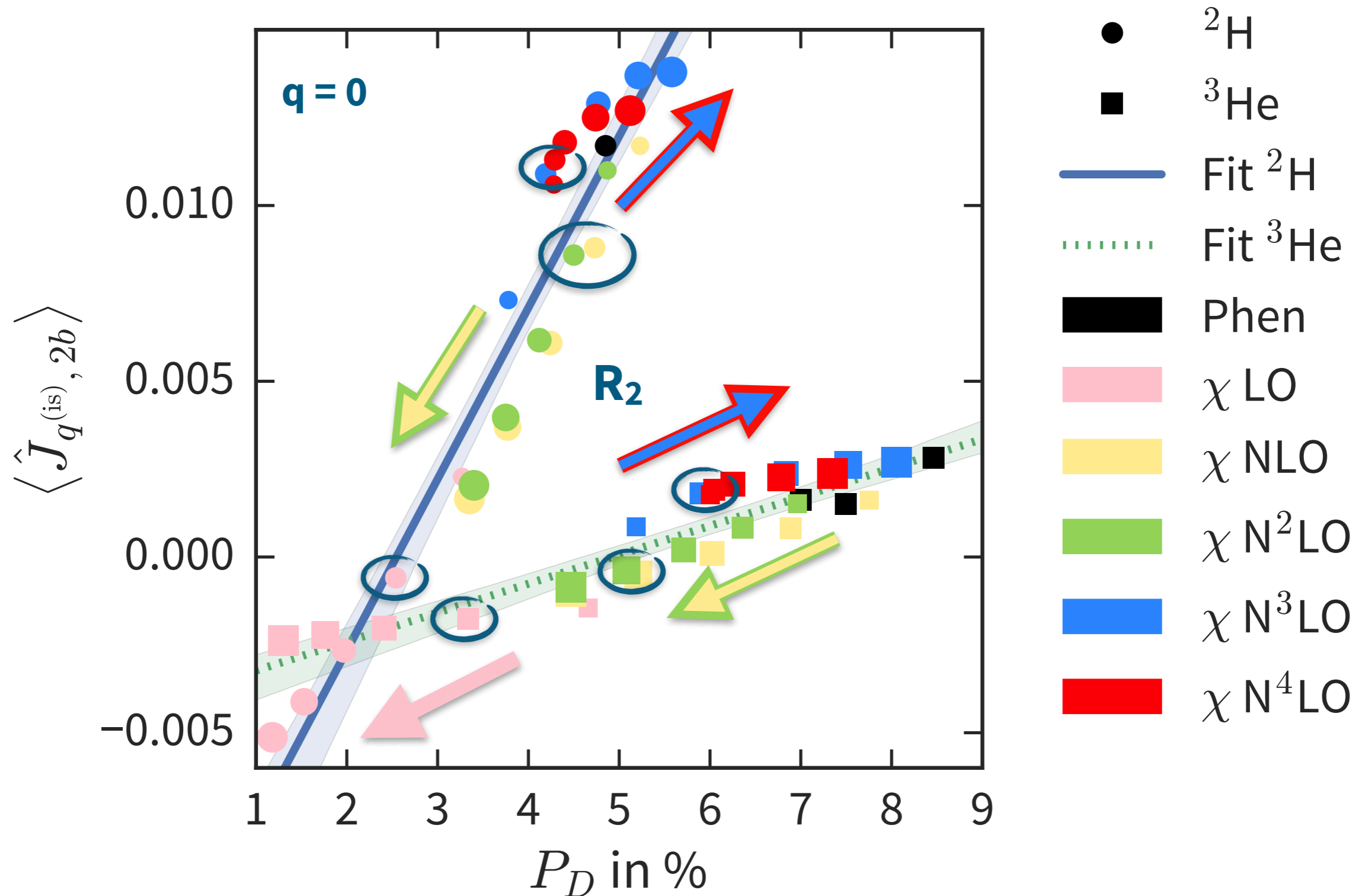
Summary Plot – Results for 2nd cut-off



Uncertainty of Two-Body Results

$$\langle \hat{J} \rangle = \langle S | \hat{J} | S \rangle + 2 \langle S | \hat{J} | D \rangle + \dots$$

- Size of marker corresponds size of cut-off
- ➔ Smallest marker = R_1 ➔ largest marker = R_5



Fine-Tuning Scenarios

The Major Difference

$$\mathcal{F}(\vec{q}^2) = (\alpha_{q^{(is)}} + \alpha_G) \mathcal{F}_{q^{(is)}}^{(0)}(\vec{q}^2) + \alpha_{q^{(is)}} \left(\mathcal{F}_{q^{(is)}, r}^{(1)}(\vec{q}^2) + \mathcal{F}_{q^{(is)}, 2b}^{(1)}(\vec{q}^2) \right) + \alpha_G \mathcal{F}_{G, 2b}^{(3)} + \dots$$

N²LO Waves

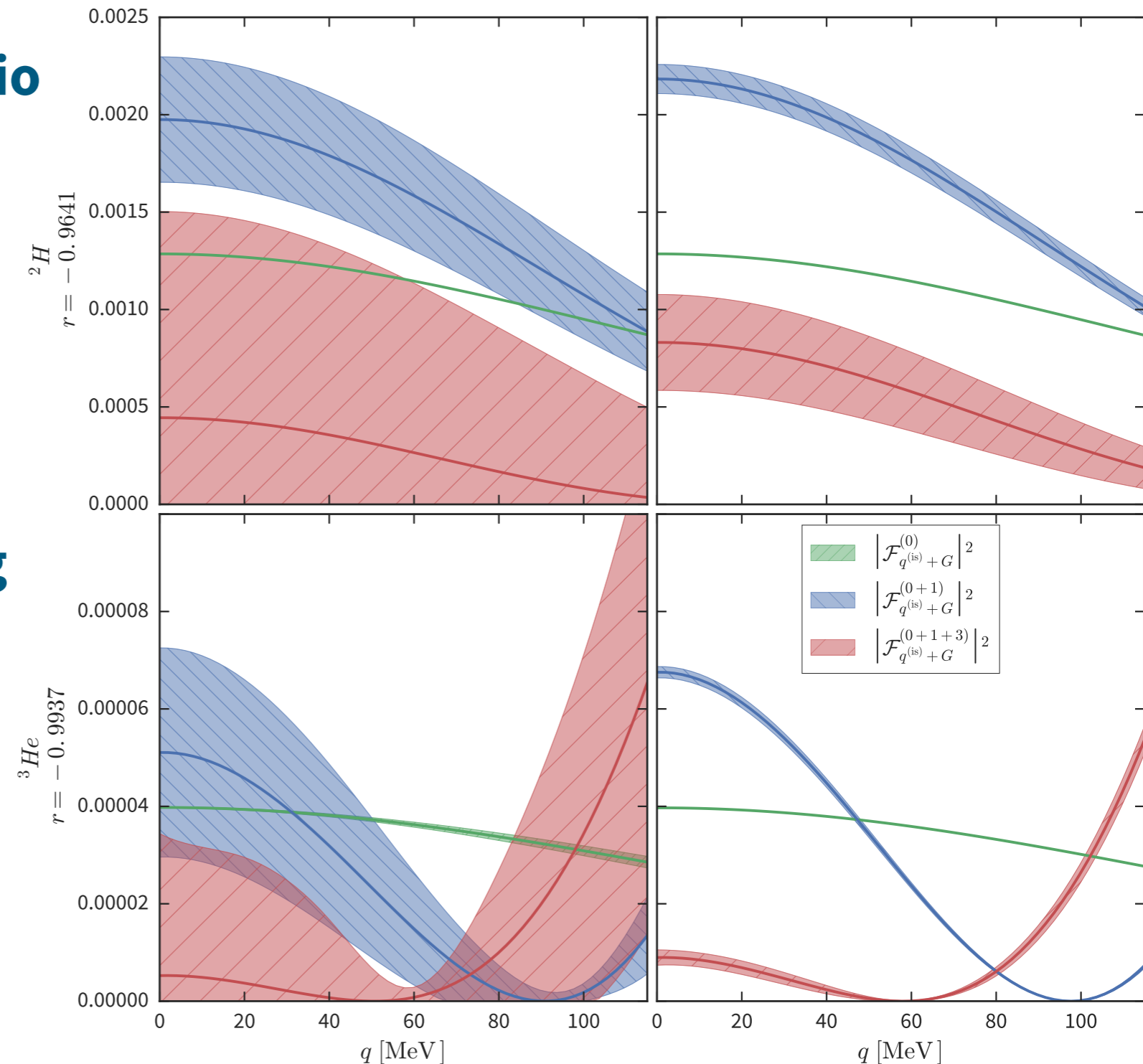
N³LO Waves

Simulate Fine-Tuning Scenario

- Define $r = \frac{\alpha_G}{\alpha_{q^{(is)}}}$
- Choose 'r' such that cross section gets 50% correction from NLO currents (for AV18)

Consequences of Fine-Tuning

- Current power counting mixed up
- Seemingly distinguishable results only for N³LO waves



Intentions for the Future

(Drawings are not in scale)

Soon...

...a few months...

...in the future

